## Page 1

# Lecture #8: Quantum Mechanical Harmonic Oscillator

Last time

Classical Mechanical Harmonic Oscillator

\* angular frequency 
$$\omega = [k/\mu]^{1/2}$$

\* 
$$\mu = \frac{m_1 m_2}{m_1 + m_2}$$
 reduced mass

From 
$$F = ma$$
 we get  $\frac{d^2x}{dt^2} = -\frac{k}{m}x$  [we get  $x(t)$  from this, not from  $\psi(x)$ ]

 $x(t) = A\sin\omega t + B\cos\omega t = C\sin(\omega t + \phi)$ 

get A, B or  $C, \phi$  from initial conditions of pluck

turning point 
$$x_{\pm}(E) = \pm \left(\frac{2E}{k}\right)^{1/2}$$
 from  $E = V(x_{\pm}(E))$ 

 $v,\omega,\tau$   $T(t),\overline{T}$  (kinetic energy)  $V(t),\overline{V}$  (potential energy)

<u>Today</u>

- \* simplify Schrödinger Equation to get rid of constant factors
- \* solution: Gaussian envelope × Hermite polynomials
- \* pictures
- \* semiclassical interpretation (not in most texts): combination of classical mechanics with  $\lambda(x) = h/p(x)$  (a unique source of insight)
- \* vibrational transition intensities and "selection rules"

Quantum Mechanical Harmonic Oscillator (McQuarrie, Chapters 5.5, 5.8-10)

$$\widehat{H} = \widehat{T} + \widehat{V} = \frac{\widehat{p}^2}{2\mu} + \frac{1}{2}k\widehat{x}^2$$
$$= -\frac{\hbar^2}{2\mu}\frac{\partial}{\partial x^2} + \frac{1}{2}k\widehat{x}^2$$

We can "clean up" this equation by making the substitution

 $\xi = \alpha^{1/2} x$  where  $\alpha = (k\mu)^{1/2} / \hbar$  ( $\xi$  is dimensionless,

which makes the equation "universal")

$$\frac{\hbar^2}{2\mu} \frac{\partial^2}{\partial x^2} = \frac{\hbar^2}{2\mu} \alpha \frac{\partial^2}{\partial \xi^2} = \frac{\hbar}{2} \left(\frac{k}{\mu}\right)^{1/2} \frac{\partial^2}{\partial \xi^2} = \frac{\hbar}{2} \omega \frac{\partial^2}{\partial \xi^2}$$

$$\begin{bmatrix} \text{because } \frac{\partial}{\partial x} = \frac{\partial}{\partial \xi} \frac{\partial \xi}{\partial x} \\ \frac{\partial}{\partial x} = \frac{\partial}{\partial \xi} \frac{\partial \xi}{\partial x} \end{bmatrix}$$

$$\frac{1}{2} kx^2 = \frac{1}{2} k \left(\frac{1}{\alpha}\right) \xi^2$$

$$= \frac{1}{2} \left(\frac{k}{\mu}\right)^{1/2} \hbar \xi^2 = \frac{1}{2} \hbar \omega \xi^2$$

 $\widehat{H} = \frac{\hbar\omega}{2} \left[ -\frac{\partial^2}{\partial\xi^2} + \xi^2 \right] \qquad (\text{ much simpler form})$  $\widehat{H}\psi(\xi) = E\psi(\xi)$  $0 = \left[ -\frac{\partial^2}{\partial\xi^2} + \xi^2 - \frac{2E}{\hbar\omega} \right] \psi \qquad (\text{entire differential equation, except} \\ \psi, \text{ is dimensionless})$ 

One can convert this into the Hermite differential equation by making the substitution  $\psi(\xi) = e^{-\xi^2/2} f(\xi)$  and finding a new differential equation for  $f(\xi)$ . The reason for doing this is that  $e^{-\xi^2/2}$  ensures that  $\psi \to 0$  as  $|\xi| \to \infty$ . Note that letting  $\xi^2 \to \infty$  means that  $2E/\hbar\omega$  is negligible with respect to  $\xi^2$ . What is the solution to the differential equation if we ignore the  $\frac{2E}{\hbar\omega}$  term? **Gaussian** 

This is a very clean form of the Schrödinger equation because all of the  $k,\mu$ -specific factors are absorbed into a dimensionless  $\xi$  variable. Why would we want this?

The Hermite polynomials (in integer powers of  $\xi$ ) are solutions to the differential equation

$$\frac{d^2 H_n}{d\xi^2} - 2\xi \frac{dH_n}{d\xi} + 2nH_n = 0.$$
 (Hermite equation)

There are very convenient "recursion relations" that relate the  $H_{n-1}$  to the  $H_n$ , etc.

$$\frac{dH_n}{d\xi} = 2nH_{n-1}(\xi) \ .$$

This looks like the effect of  $\hat{p}$  on  $H_n$  (because the leading term in  $H_n$  is  $\xi^n$ ).

$$H_{n+1}(\xi) = 2\xi H_n(\xi) - 2nH_{n-1}(\xi), \text{ rearranging}$$
  
$$\xi H_n(\xi) = \frac{1}{2}H_{n+1}(\xi) + nH_{n-1}(\xi).$$

This looks like the effect of  $\hat{x}$  on  $H_n$ . We will use this second recursion relation to compute integrals of the form

$$\int d\xi \, \psi_n^* \xi^m \psi_p \, . \, (n, m, p \text{ are integers}).$$

These recursion relations enable us to evaluate all integrals of the form

$$\int \Psi_{\nu}(\hat{x}^n \hat{p}^m) \Psi_{\nu+\ell} dx$$

(we will postpone the actual evaluation until next lecture) and we derive the "selection rule" for nonzero integrals

$$\ell = n + m, n + m - 2, \dots - (n + m)$$

There is also a general expression (Rodrigues formula) for the  $H_n$ 

$$H_n(\xi) = (-1)^n e^{\xi^2} \frac{d^n}{d\xi^n} e^{-\xi^2}$$

The Hermite equation is a well known (to mathematicians) differential equation.

The solutions of the Schrödinger equation are



What do we know about orthogonality? Based on results derivable from postulates? Non-degenerate eigenvalues.

 $\int dx \psi_v^* \psi_{v'} = \delta_{vv'} \qquad \text{orthonormal}$ 

Semi-Classical Picture – applicable to all 1-dimensional problems for insight, prediction, and obtaining  $\psi(x)$  without solving any differential equation.

Classical:  $T(x) = E - V(x) = p(x)^2/2\mu$  $p(x) = [2m(E - V(x))]^{1/2}$ 

p(x) is momentum in classical mechanics but just a convenient function in quantum mechanics.



Quantum: de Broglie  $\lambda = \frac{h}{p}$ valid not just for free particle or a piecewise constant potential

$$(=)$$

Semi-classical:

$$\lambda(x) \equiv \frac{h}{p_{\text{classical}}(x)}$$

\* pair of nodes nearest to *x* are spaced by  $\lambda(x)/2$ .

Qualitative Shapes of  $\psi_v(x)$ :

- \* exponentially damped envelope, extending into non-classical regions (!!!!)
- \* oscillations within classically allowed region with number of internal nodes equal to the quantum number
- \* even *v*, even function, antinode at x = 0
- \* outer lobes (near  $x_+$  and  $x_-$ ) are largest [see McQuarrie, page 226, Fig. 5.10 right side]
- \* envelope within classically allowed region resembles what you expect from classical mechanics

 $\Psi^*(x)\Psi(x)dx \propto \frac{dx}{v_{\text{classical}}} \qquad (\text{slow speed} \leftrightarrow \text{high probability})$  $v_{\text{classical}} = p_{\text{classical}} / \mu = \frac{1}{\mu} \left\{ 2\mu \left[ E - V(x) \right] \right\}^{1/2}$ 

To get the proportionality constant, consider the fraction of time the particle is found between x and x + dx:

$$\frac{\text{time } (x \to x + dx)}{\text{time } (x_{-} \text{ to } x_{+})} = \begin{cases} \text{ probability of finding particle moving} \\ \text{ to right between } x \text{ and } x + dx \end{cases} = \frac{dx/v_{\text{classical}}(x)}{\tau/2}$$

$$\tau = \frac{1}{\nu} = \frac{2\pi}{\omega} = 2\pi \left(\frac{\mu}{k}\right)^{1/2} \left[\nu = \omega/2\pi, \omega = \left[k/\mu\right]^{1/2}\right]$$
$$\psi^*(x)\psi(x)dx = \frac{dx}{\left(\nu_{\text{classical}}\right)\left(\tau/2\right)} = \frac{dx}{\left\{\frac{2}{\mu}\left[E - V(t)\right]\right\}^{1/2}} \frac{1}{\pi} \left(\frac{\mu}{k}\right)^{-1/2} = \left[\frac{k/2\pi^2}{E - V(x)}\right]^{1/2} dx$$
$$\psi^*(x)\psi(x) = \left[\frac{k/2\pi^2}{E - V(x)}\right]^{1/2} \text{ gives the classical average of } \psi^*(x)\psi(x) \text{ near } x$$

(but not the phase). [To get the classical envelope, assume that the maximum value of  $\psi^*\psi$  is twice the average value. [This is always a good approximation for a rapidly oscillating always positive function.] Thus the envelope of

$$\Psi^*(x)\Psi(x)$$
 envelope is  $\left[\frac{2k/\pi^2}{E-V(x)}\right]^{1/2}$ 

- \* node spacing
  - (1) recall  $\lambda = \frac{h}{p(x)}$ , use classical ideas to *qualitatively* locate nodes,
  - (2) nodes are closest together when *p* is largest (near x = 0), envelope and node spacing allows you to sketch  $\psi^*(x)\psi(x)$  without solving a differential equation.
  - (3) or compute a "phase integral"  $\frac{\Delta x}{\lambda(x)} = \frac{1}{2}$  (want to find value of  $\Delta x$  that is

equal to  $\lambda/2$ , the distance between nodes)

replace  $\lambda(x)$  by h/p(x). We get

$$p(x)\Delta x = \frac{h}{2}$$
 as the distance,  $\Delta x$ , between nodes  $\left[\text{more accurately}, \int_{x}^{x+\lambda/2} dx p(x) = \frac{h}{2}\right]$ 

phase integral  $\frac{2}{h} \int_{x_1}^{x_2} p_E(x) dx$  tells us *how many nodes* there are between  $x_1$  and  $x_2$  at energy *E*. This is the same as knowing how many bound energy levels lie at or below *E*.





#### FIGURE 5.10

(a) The normalized harmonic-oscillator wave functions. (b) The probability densities for a harmonic oscillator. As in Figure 5.7, the potential energy is indicated by the parabolas in (a) and (b).

T A B L E 5.3 The First Few Hermite Polynomials<sup>a</sup>

$H_0(\xi) = 1$	$H_1(\xi) = 2\xi$
$H_2(\xi) = 4\xi^2 - 2$	$H_3(\xi) = 8\xi^3 - 12\xi$
$H_4(\xi) = 16\xi^4 - 48\xi^2 + 12$	$H_5(\xi) = 32\xi^5 - 160\xi^3 + 120\xi$
$H_2(\xi) = 4\xi^2 - 2$ $H_4(\xi) = 16\xi^4 - 48\xi^2 + 12$	$H_3(\xi) = 8\xi^3 - 12\xi$ $H_5(\xi) = 32\xi^5 - 160\xi^3 + 120\xi$

a. The variable  $\xi$  is equal to  $\alpha^{1/2}x$ , where  $\alpha = (k\mu)^{1/2}/\hbar$ .

#### TABLE 5.4

The First Few Harmonic-Oscillator Wave Functions, Equation 5.44 a

$$\begin{split} \psi_0(x) &= |0\rangle = \left(\frac{\alpha}{\pi}\right)^{1/4} e^{-\alpha x^2/2} & \psi_2(x) = |2\rangle = \left(\frac{\alpha}{4\pi}\right)^{1/4} (2\alpha x^2 - 1)e^{-\alpha x^2/2} \\ \psi_1(x) &= |1\rangle = \left(\frac{4\alpha^3}{\pi}\right)^{1/4} x e^{-\alpha x^2/2} & \psi_3(x) = |3\rangle = \left(\frac{\alpha^3}{9\pi}\right)^{1/4} (2\alpha x^3 - 3x)e^{-\alpha x^2/2} \end{split}$$

a. The parameter  $\alpha = (k\mu)^{1/2}/\hbar$ .

McQuarrie, Donald A. Quantum Chemistry. 2nd ed. Sausalito, CA: University Science Books, 2007. Used with permission.

- \* non-classical tails [(extend into region where E < V(x)]
- \*  $x_{+}$  lobe positive by convention
- \* lobes nearest  $x_{-}$  and  $x_{+}$  largest
- \* nodes closest together near x = 0
- \* no zero crossings in classically forbidden region

What about pictures of  $\psi^*(x)\psi(x)$ ?

## Non-Lecture

What do we do with these HO wavefunctions?

1. calculate relative intensities of vibrational transitions

2. Use perturbation theory (Lectures #14, #15 and #18) to compute consequences of higher than quadratic terms in V(x)

# e.g. for Morse oscillator

$$\frac{E_{v}}{hc} = \tilde{\omega} (v + 1/2) - \tilde{\omega} v + 1/2)^{2}$$
 (~ means cm<sup>-1</sup> units)  
"anharmonicity", comes mostly  
from x<sup>3</sup> and x<sup>4</sup> terms in V(x)

Spectral intensities

$$I_{f_i} \propto \left| \int dx \psi_{v_f}^* \mu(x) \psi_{v_i} \right|^2$$

$$\stackrel{\text{electric dipole moment}}{\underset{(\text{HCl vs. H}_2, N_2, O_2, Cl_2)}{\underset{(\text{think of radio antenna)}}}} \int \frac{1}{|v_i|^{v_f}}$$



Two contributions to vibrational overtone transitions

- \* mechanical anharmonicity (Morse potential)
- \* electronic anharmonicity (higher derivatives of  $\mu(x)$ )

Page 9

In the following we will look only at the electronic anharmonicity contributions.

$$\int dx \psi_{v_f}^* \left[ \mu_0 + \mu_1 x + \frac{1}{2} \mu_2 x^2 \right] \psi_{v_i} = \mu_0 \int dx \underbrace{\psi_{v_f}^* \psi_{v_i}}_{\substack{\text{ortho-normal} \\ \delta_{v_f v_i}}} + \mu_1 \underbrace{\int dx \, \psi_{v_f}^* x \psi_{v_i}}_{\substack{\text{see recursion} \\ \text{relationship}}} + \frac{\mu_2}{2} \int dx \, \psi_{v_f}^* x^2 \psi_{v_i}$$

$$H_{n+1}(\xi) = 2\xi H_n(\xi) - 2nH_{n-1}(\xi)$$
  
$$\underbrace{\xi H_n(\xi)}_{x\psi_v} = \frac{1}{2}\underbrace{H_{n+1}}_{\psi_{v+1}}(\xi) + n\underbrace{H_{n-1}}_{\psi_{v-1}}(\xi)$$

selection rule:  $\Delta v = \pm 1$ 

for  $x^2$  term (evaluate in two steps)

$$\xi^{2}H_{n} = \frac{1}{2}\xi H_{n+1} + n\xi H_{n-1}$$
  
=  $\frac{1}{2}\left(\frac{1}{2}H_{n+2} + (n+1)H_{n}\right) + n\left(\frac{1}{2}H_{n} + (n-1)H_{n-2}\right)$   
 $\Delta v = 0, \pm 2$ 

Next time:  $\mathbf{a}^{\dagger}$ ,  $\mathbf{a}$  treatment

5.61 Physical Chemistry Fall 2013

For information about citing these materials or our Terms of Use, visit: http://ocw.mit.edu/terms.