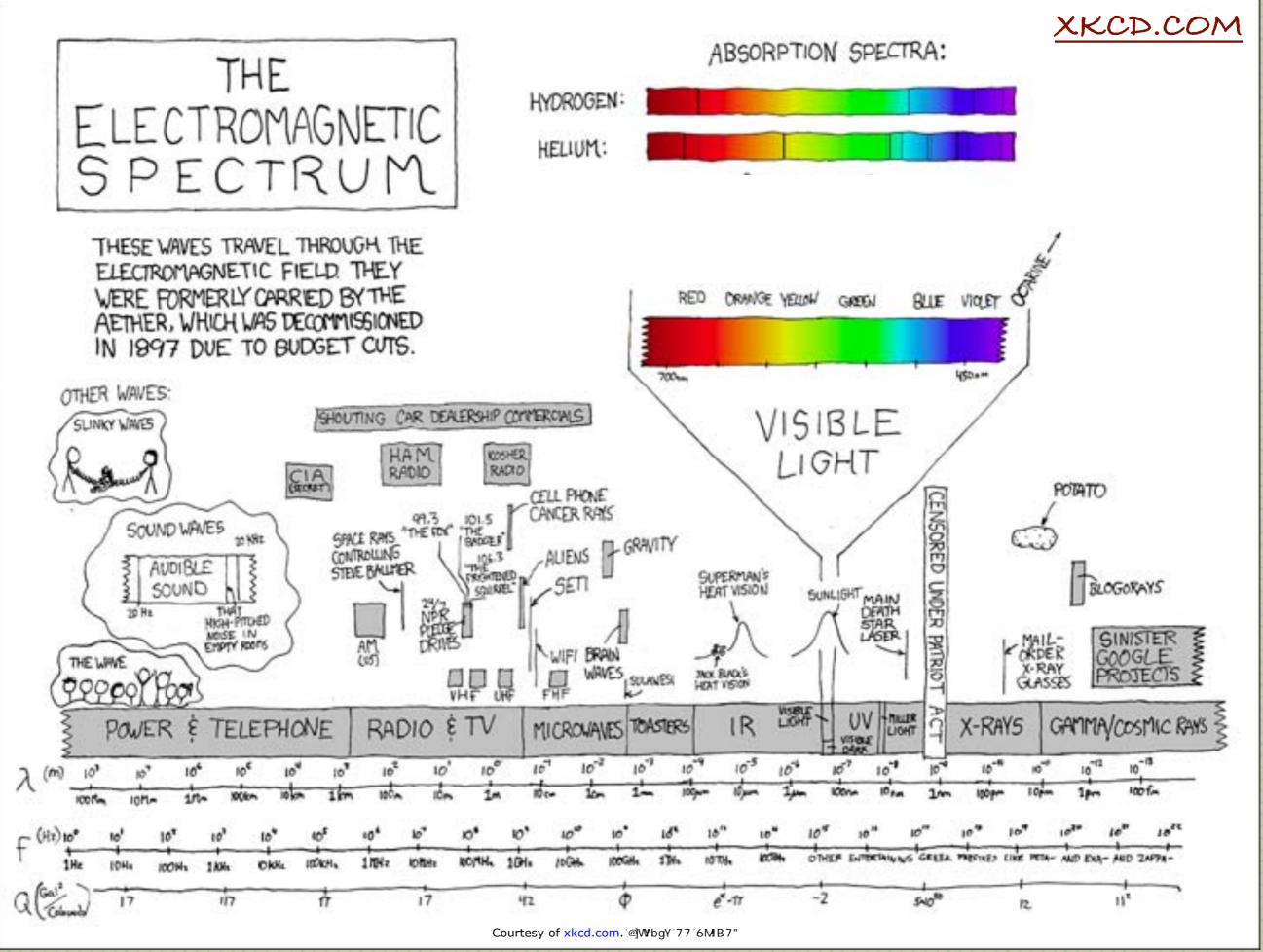
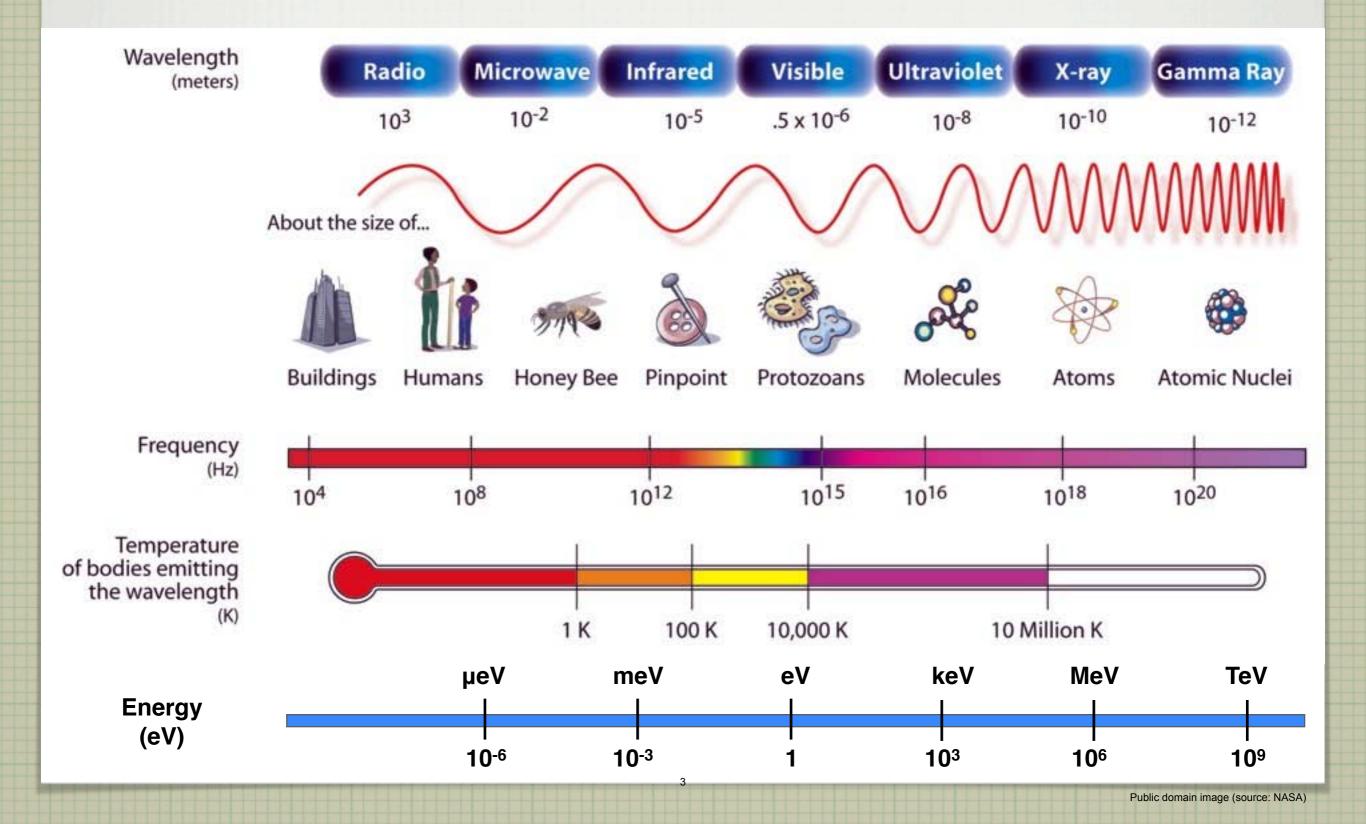
GAMMA DECAY



ELECTROMAGNETIC SPECTRUM



DIPOLE RADIATION

□ Rate from Fermi's Golden Rule + Density of states:

$$W = \frac{2\pi}{\hbar} |\langle \psi_f | \hat{V} | \psi_i \rangle|^2 \rho(E_f) = \frac{\omega^3}{2\pi c^3 \hbar} |\langle \hat{r} \rangle|^2 \sin^2 \theta \, d\Omega$$

□ Integrating over angles:

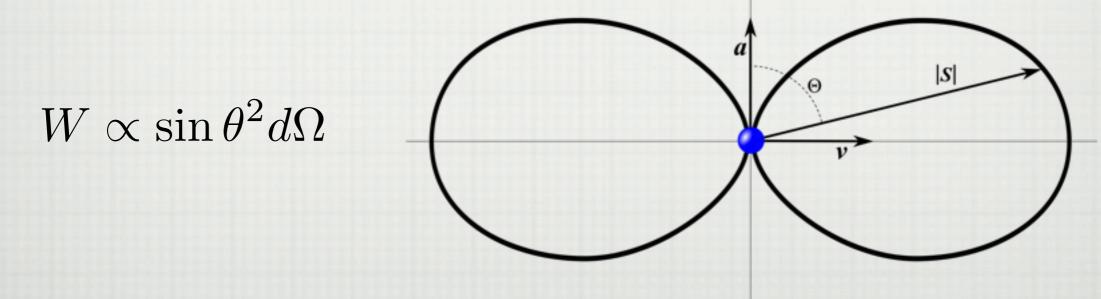
$$\lambda(E1) = \frac{4}{3} \frac{e^2 \omega^3}{\hbar c^3} |\langle r \rangle|^2$$

DIPOLE APPROXIMATION

In deriving the dipole emission formula we only kept lowest order expansion:

$$\vec{A} \propto a_k^{\dagger} e^{i\vec{k}\cdot\vec{r}}\vec{\epsilon}_k \approx a_k^{\dagger}(1+i\vec{k}\cdot\vec{r}+\dots)\vec{\epsilon}_k \to a_k^{\dagger}\vec{\epsilon}_k$$

□ This yields the typical dipole emission pattern:



In QM the angular distribution is related to the photon angular momentum

BEYOND THE DIPOLE

- Higher order terms in the expansion give rise to gamma emission
 - □ with different angular-dependence pattern
 - and higher angular momentum for the gamma photon emitted

$$\vec{A} \propto a_k^{\dagger} e^{i\vec{k}\cdot\vec{r}} \vec{\epsilon}_k \approx a_k^{\dagger} \sum_{\ell} \frac{(i\vec{k}\cdot\vec{r})^{\ell}}{\ell!} \vec{\epsilon}_k$$

 \Box Each ℓ term contributes to a different decay rate.

MULTIPOLE RADIATION

Electric multipole

$$\lambda(E\ell) = \frac{8\pi(\ell+1)}{\ell[(2\ell+1)!!]^2} \frac{e^2}{\hbar c} \left(\frac{E}{\hbar c}\right)^{2\ell+1} \left(\frac{3}{\ell+3}\right)^2 c \left\langle |\hat{\vec{r}}| \right\rangle^{2\ell}$$

$$\left\langle \left| \hat{\vec{r}} \right| \right\rangle \approx R_0 A^{1/3}$$

Rates:

 $\lambda(E1) = 1.0 \times 10^{14} A^{2/3} E^3$ $\lambda(E2) = 7.3 \times 10^7 A^{4/3} E^5$ $\lambda(E3) = 34A^2E^7$ $\lambda(E4) = 1.1 \times 10^{-5} A^{8/3} E^9$

MULTIPOLE RADIATION

□ Magnetic multipole

$$\lambda(M\ell) = \frac{8\pi(\ell+1)}{\ell[(2\ell+1)!!]^2} \frac{e^2}{\hbar c} \frac{E}{\hbar c}^{2\ell+1} \left(\frac{3}{\ell+3}\right)^2 c \left\langle |\hat{\vec{r}}| \right\rangle^{2\ell-2} \left[\frac{\hbar}{m_p c} \left(\mu_p - \frac{1}{\ell+1}\right)\right]$$

□ Rates: $\lambda(M1) = 5.6 \times 10^{13} E^3$

 $\lambda(M2) = 3.5 \times 10^7 A^{2/3} E^5$ $\lambda(M3) = 16A^{4/3} E^7$ $\lambda(M4) = 4.5 \times 10^{-6} A^2 E^9$

WHICH TRANSITION?

- □ The lowest multipole dominates:
 - □ Lower multipoles decay faster (higher rates)
 - □ Electric multipoles are faster than magnetic multipoles

Why don't we always only observe electric dipole (E1) radiation?

SELECTION RULES

- \Box The multipole ℓ is related to the gamma angular momentum
 - □ the angular momentum must be conserved in gamma decay
- $\square \text{ Possible } \ell: \quad |I_f I_i| \le \ell_{\gamma} \le I_f + I_i$
- □ Parity: $(-1)^{\ell}$ for Electric and $(-1)^{\ell-1}$ for Magnetic: parity must be conserved, $\Pi_{\gamma} = \Pi_{i} \Pi_{f}$

Multipolarity	Angular	Parity	Multipolarity	Angular	Parity
	Momentum l	Π		Momentum l	Π
M1	1	+	E1	1	-
M2	2	-	E2	2	+
M3	3	+	E3	3	-
M4	4	-	E4	4	+
M5	5	+	E5	5	-

WHICH TRANSITION?

- □ The lowest permitted multipole dominates
- Electric multipoles are more probable than the same magnetic multipole by a factor 100 $\lambda(El)$

$$\frac{\lambda(El)}{\lambda(Ml)} \approx 10^2$$

Emission from the multipole *l*+1 is 10⁻⁵ times less probable than the *l*-multipole emission

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$$\frac{\lambda(E,l+1)}{\lambda(El)} \approx 10^{-5}, \qquad \frac{\lambda(M,l+1)}{\lambda(Ml)} \approx 10^{-5}$$

WHICH TRANSITION?

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□ Combining the two rules:

$$\frac{\lambda(E, l+1)}{\lambda(Ml)} \approx 10^{-3},$$

$$\frac{\lambda(M,l+1)}{\lambda(El)} \approx 10^{-7}$$

□ Thus E2 competes with M1

□ But M2 does not compete with E1

INTERNAL CONVERSION

In some cases energy is not released in the form of gamma photons, but carried away by an electron:

 ${}^{A}_{Z}X^{*} \rightarrow {}^{A}_{Z}X^{+} + e^{-}$

This process is called Internal Conversion

It is the only process possible, when selection rules do not allow any of the multipole transitions:

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- □ e.g. even-even nuclides, decay from a 0⁺ level
- □ the photon cannot have zero angular momentum.

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