8.821 String Theory Fall 2008

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## 8.821 F2008 Lecture 13: Masses of fields and dimensions of operators

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In today's lecture we will talk about:

- 1. AdS wave equation near the boundary.
- 2. Masses and operator dimensions:  $\Delta(\Delta D) = m^2 L^2$ .

<u>Erratum</u>: The massive geodesic equation  $\ddot{x} + \Gamma \dot{x} \dot{x} = 0$  assumes that the dot differentiates with respect to proper time.

<u>Recap</u>: Consider a scalar in  $AdS_{p+2}$  (where p+1 is the number of spacetime dimensions that the field theory lives in). Let the metric be:

$$ds^{2} = L^{2} \frac{dz^{2} + dx^{\mu} dx_{\mu}}{z^{2}},$$
(1)

then the action takes the form:

$$S[\phi] = -\frac{\kappa}{2} \int d^{p+1}x \sqrt{g} \left( (\partial \phi)^2 + m^2 \phi^2 + b\phi^3 + \dots \right),$$
(2)

where  $(\partial \phi)^2 \equiv g^{AB} \partial_A \phi \partial_B \phi$  and  $x^A = (z, x^{\mu})$ . Our goal is to evaluate:

$$\ln \langle \exp^{-\int d^D x \, \phi_0 \, O} \rangle_{CFT} = \operatorname{extremum}_{[\phi \mid \phi \to \phi_0 \ at \ z = \epsilon]} S[\phi], \tag{3}$$

where  $S[\phi] \equiv S[\phi^*(\phi_0)] \equiv W[\phi_0]$ , i.e. by using the solution to the equation of motion subject to boundary conditions. Now Taylor expand:

$$W[\phi_0] = W[0] + \int d^D x \ \phi_0(x) G_1(x) + \frac{1}{2} \int \int d^D x_1 d^D x_2 \ \phi_0(x_1) \phi_0(x_2) G_2(x_1, x_2) + \dots$$
(4)

where

$$G_1(x) = \langle O(x) \rangle = \frac{\delta W}{\delta \phi_0(x)} |_{\phi_0=0}, \tag{5}$$

$$G_2(x) = \langle O(x_1)O(x_2) \rangle_c = \frac{\delta^2 W}{\delta \phi_0(x_1)\delta \phi_0(x_2)}|_{\phi_0=0}.$$
 (6)

Now if there is no instability, then  $\phi_0$  is small and so is  $\phi$ , so you can ignore third order terms in  $\phi$ . From last time:

$$S[\phi] = \frac{\kappa}{2} \int_{AdS_{p+2}} d^{p+2}x \sqrt{g} \left[\phi \left(-\nabla^2 + m^2\right)\phi + \mathcal{O}(\phi^3)\right] - \frac{\kappa}{2} \int_{\partial AdS} d^{p+1}x \sqrt{\gamma} \phi \left(n.\partial\right)\phi, \tag{7}$$

where the last term is the boundary action, n is a normalized vector perpendicular to the boundary and

$$\nabla^2 = \frac{1}{\sqrt{g}} \partial_A(\sqrt{g} g^{AB} \partial_B). \tag{8}$$

Now if the scalar field satisfies the wave equation:

$$(-\nabla^2 + m^2)\phi^* = 0, (9)$$

$$W[\phi_0] = S_{bdy}[\phi^*[\phi_0]], \tag{10}$$

then we can use translational invariance in p + 1 dimensions,  $x^{\mu} \to x^{\mu} + a^{\mu}$ , in order to Fourier decompose the scalar field:

$$\phi(z, x^{\mu}) = e^{ik \cdot x} f_k(z). \tag{11}$$

Now, substituting (11) into (9) and assuming that the metric only depends on z we get:

$$0 = (g^{\mu\nu}k_{\mu}k_{\nu} - \frac{1}{\sqrt{g}}\partial_{z}(\sqrt{g}g^{zz}\partial_{z}) + m^{2})f_{k}(z)$$
(12)

$$= \frac{1}{L^2} [z^2 k^2 - z^{D+1} \partial_z (z^{-D+1} \partial_z) + m^2 L^2] f_k, \qquad (13)$$

where we have used  $g^{\mu\nu} = (z/L)^2 \delta^{\mu\nu}$ . The solutions of (12) are Bessel functions but we can learn a lot without using their full form. For example, look at the solutions near the boundary (i.e.  $z \to 0$ ). In this limit we have power law solutions, which are spoiled by the  $z^2k^2$  term. Try using  $f_k = z^{\Delta}$  in (12):

$$0 = k^2 z^{2+\Delta} - z^{D+1} \partial_z (\Delta z^{-D+\Delta}) + m^2 L^2 z^{\Delta}$$

$$\tag{14}$$

$$= (k^{2}z^{2} - \Delta(\Delta - D) + m^{2}L^{2})z^{\Delta}, \qquad (15)$$

and for  $z \to 0$  we get:

$$\Delta(\Delta - D) = m^2 L^2 \tag{16}$$

The two roots for (16) are

$$\Delta_{\pm} = \frac{D}{2} \pm \sqrt{\left(\frac{D}{2}\right)^2 + m^2 L^2}.$$
(17)

## Comments

- The solution proportional to  $z^{\Delta_{-}}$  is bigger near  $z \to 0$ .
- $\Delta_+ > 0 \ \forall \ m$ , therefore  $z^{\Delta_+}$  decays near the boundary.
- $\Delta_+ + \Delta_- = D$ .

Next, we want to improve the boundary conditions that allow solutions, so take:

$$\phi(x,z)|_{z=\epsilon} = \phi_0(x,\epsilon) = \epsilon^{\Delta_-} \phi_0^{Ren}(x), \tag{18}$$

where  $\phi_0^{Ren}$  is the renormalized field. Now with this boundary condition,  $\phi(z, x)$  is finite when  $\epsilon \to 0$ , since  $\phi_0^{Ren}$  is finite in this limit.

Wavefunction renormalization of O (Heuristic but useful)

Suppose:

$$S_{bdy} \quad \ni \quad \int_{z=\epsilon} d^{p+1}x \,\sqrt{\gamma_{\epsilon}} \,\phi_0(x,\epsilon) \,O(x,\epsilon) \tag{19}$$

$$= \int d^{D}x \, \left(\frac{L}{\epsilon}\right)^{D} \left(\epsilon^{\Delta_{-}} \phi_{0}^{Ren}(x)\right) O(x,\epsilon), \tag{20}$$

where we have used  $\sqrt{\gamma} = (L/\epsilon)^D$ . Demanding this to be finite as  $\epsilon \to 0$  we get:

$$O(x,\epsilon) \sim \epsilon^{D-\Delta_{-}} O^{Ren}(x)$$
 (21)

$$= \epsilon^{\Delta_+} O^{Ren}(x), \tag{22}$$

where in the last line we have used  $\Delta_+ + \Delta_- = D$ . Therefore, the scaling of  $O^{Ren}$  is  $\Delta_+ \equiv \Delta$ .

## Comments

- We will soon see that  $\langle O(x)O(0)\rangle \sim \frac{1}{|x|^{2\Delta}}$ .
- We had a second order ODE, therefore we need two conditions in order to determine a solution (for each k). So far we have imposed:
  - 1. For  $z \to \epsilon$ ,  $\phi \sim z^{\Delta_-} \phi_0 + (\text{terms subleading in } z)$ . Now we will also impose
  - 2.  $\phi$  regular in the interior of AdS (i.e. at  $z \to \infty$ ).

## Comments on $\Delta$

1. The  $\epsilon^{\Delta_{-}}$  factor is independent of k and x, which is a consequence of a local QFT (this fails in exotic examples).

2. <u>Relevantness</u>: Since  $m^2 > 0 \implies \Delta \equiv \Delta_+ > D$ , so  $O_{\Delta}$  is an irrelevant operator. This means that if you perturb the CFT by adding  $O_{\Delta}$  to the Lagrangian, then:

$$\Delta S = \int d^D x \,(\text{mass})^{D-\Delta} O_{\Delta},\tag{23}$$

where the exponent is negative, so the effects of such an operator go away in the IR. For example, consider a dilaton mode with l > 0, its mass is given by (for D = 4):

$$m^2 = \frac{(l+4)l}{L^2}.$$
 (24)

The operator corresponding to this is:

$$\operatorname{tr}(F^2 X^{i_1 \dots i_l}),\tag{25}$$

with  $\Delta = 4 + l > D$ , therefore it is an irrelevant operator. Now consider a dilaton mode with l = 0: then  $m^2 = 0$ , therefore,  $\Delta = D$  and hence it corresponds to a marginal operator (an example of such operator is the Lagrangian). If  $m^2 < 0$ , then  $\Delta < D$ , so it corresponds to a relevant operator, but it is ok if  $m^2$  is not too negative ("Breitenlohner - Freedmasn (BF) - allowed tachyons" with  $-|m_{BF}|^2 \equiv -(D/2L)^2 < m^2$ ).

3. Instability: This occurs when a renormalizable mode grows with time without a source. But in order to have  $S[\phi] < \infty$ , the solution must fall off at the boundary. This requires a gradient energy that  $\sim \frac{1}{L}$ . Note:

$$\Delta_{\pm} = \frac{D}{2} \pm \sqrt{\left(\frac{D}{2}\right)^2 + m^2 L^2}.$$
(26)

If:

$$m^2 L^2 < (\frac{D}{2})^2 \equiv -|m_{BF}|^2,$$
 (27)

then  $\Delta_{\pm}$  is complex, therefore we have  $\Delta_{-} = D/2$ , which is larger than the unitary bound. In this case,  $\phi \sim z^{\Delta_{-}}$  decays near the boundary (i.e. in the UV). In order to see the instability that occurs when  $m^{2}L^{2} < (\frac{D}{2})^{2}$  more explicitly, rewrite (9) as a Schrodinger equation, by writing  $\phi(z) = A(z)\psi(z)$ , where we choose A(z) in order to remove the first derivative of  $\psi(z)$ . Then, equation (9) becomes:

$$(-\partial_z^2 + V(z))\psi(z) = E\psi(z), \qquad (28)$$

where  $E = \omega^2 - k^2$ ,  $V(z) = \sigma/z^2$  and  $\sigma = m^2 L^2 - (D^2 - 1)/4$ . An instability occurs when E < 0, i.e.  $\omega^2 < 0$  and hence  $\phi \sim e^{i\omega t}\phi(z) = e^{+|\omega|t}\phi(z)$  grows with time. Now the claim is that  $V = \sigma/z^2$  has no negative energy states if  $\sigma > -1/4$ . Note that the notion of normalizability here and before are related (Pset 4):

$$||\psi||^2 = \int dz \,\psi^{\dagger}\psi < \infty, \tag{29}$$

and 
$$S[\phi] = \int dz \sqrt{g} \left( (\partial \phi)^2 + m^2 \right)$$
 (30)

4. The formula we found before (expression (16)) depends on the spin. For a j-form in AdS we have:

$$(\Delta + j)(\Delta + j - D) = m^2 L^2.$$
(31)

For example, for  $A_{\mu}$  massless we have:

$$\Delta(j^{\mu}) = D - 1 \quad \to \text{ conserved}, \tag{32}$$

for  $g_{\mu\nu}$  massless we have:

$$\Delta(T^{\mu\nu}) = D \quad \to \text{ required from CFT.}$$
(33)